

Interaction of particles in matter

- Particle lifetime τ : $N(t) = e^{-t/\tau}$
- Particles we detect ($\tau > 10^{-10}$ s, $c\tau > 0.03$ m)
- Charged particles
 - e^\pm (stable $m=0.511$ MeV)
 - μ^\pm ($c\tau = 659$ m $m=0.102$ GeV)
 - π^\pm ($c\tau = 7.8$ m $m=0.139$ GeV)
 - K^\pm ($c\tau = 3.7$ m $m=0.495$ GeV)
 - p^\pm (stable $m=0.937$ GeV)
- Photons γ
- Neutral hadrons
 - n ($c\tau = 2.7 \times 10^8$ m $m=0.938$ GeV)
 - K_L^0 ($c\tau = 15.5$ m $m=0.498$ GeV)

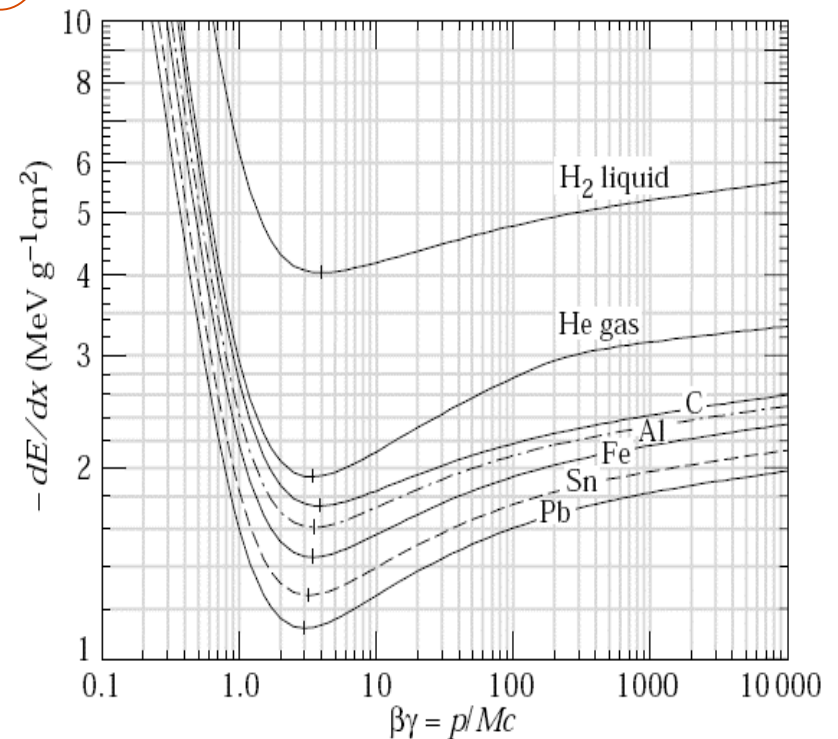
Energy loss: dE/dx

- All charged particles can lose energy via interaction with the EM fields of atoms in matter

- For $\beta\gamma < 10^3$ the dominant energy loss mechanism is via ionization of atoms:
$$-\frac{dE}{dx} = Kz^2 \frac{Z}{A} \frac{1}{\beta^2} \left[\frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 T_{\max}}{I^2} - \beta^2 - \frac{\delta}{2} \right]$$

- Note the β^{-2} dependence and the dependence on Z and A. When $\beta\gamma$ is very large the logarithm dominates.

- Minimum is at $\beta\gamma \approx 3$
- x is measured in g/cm^2 , so $x = \rho \times L$



Energy loss - radiation

- Highly relativistic particles ($\gamma > 10^3$ or so) lose energy mostly via “bremsstrahlung” radiation (emitting photons); critical energy E_c is where dE/dx loss equals radiation loss
- This radiation is emitted along the direction of motion
- Energy lost through radiation is proportional to the energy of the particle (constant fractional loss): $dE/dx|_{\text{radiation}} = E/X_0$, where X_0 is the “radiation length” and is a feature of the medium ($X_0 \sim Z^{-2}$)
- After a distance X_0 the particle retains a fraction $1/e$ of its energy: $dE_{\text{rad}}/E = dx/X_0 \rightarrow E(x) = E_{\text{initial}} \exp(-x/X_0)$
- In practice, this is always the dominant energy loss mechanism for electrons; for heavier particles dE/dx usually dominates

Electrons

- In most materials electrons lose energy predominantly by radiation above a few 10s of MeV

Fractional energy loss

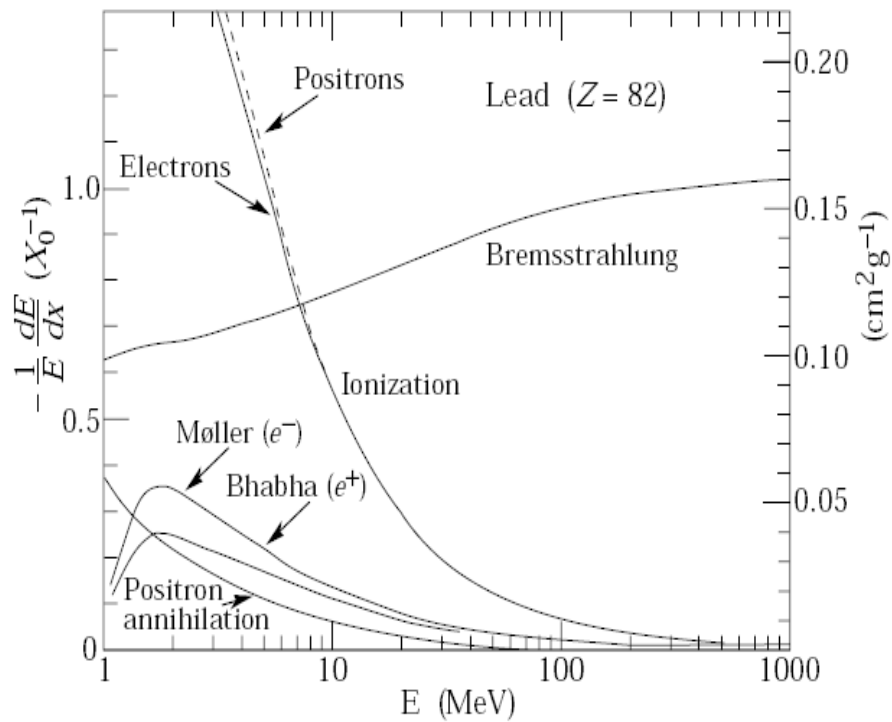
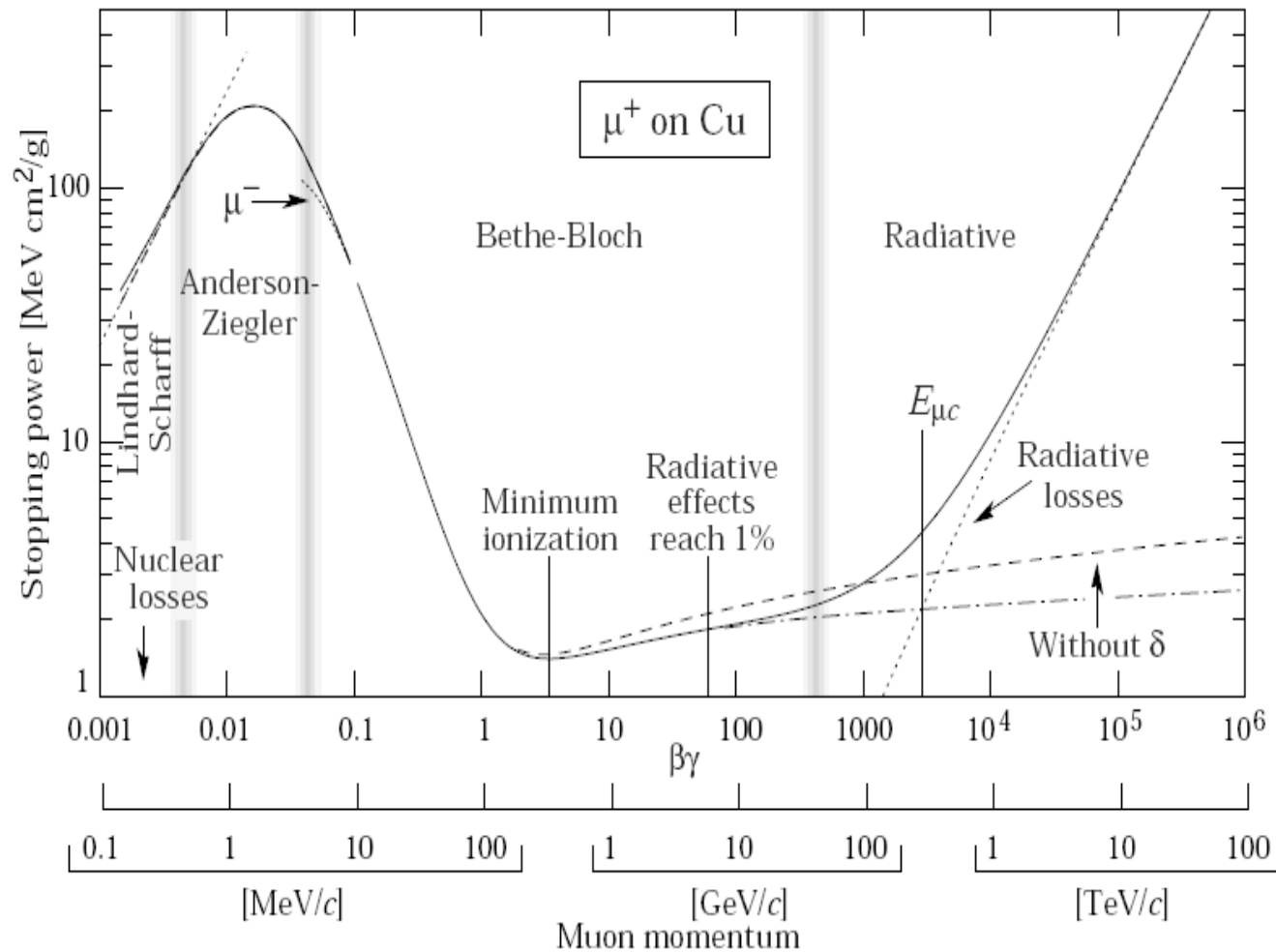


Figure 27.9: Fractional energy loss per radiation length in lead as a function of electron or positron energy. Electron (positron) scattering is considered as ionization

Muons

- Muons are like electrons, only ~ 200 times heavier.
- The mass makes a big difference: dominant energy loss is by dE/dx instead of radiation \rightarrow much longer range in matter than either electrons (which radiate photons) or hadrons (which interact strongly)
- Put a detector behind enough shielding and the particles that come through are (mostly) muons (consider, e.g., the atmosphere)

Energy loss of μ in Cu



Cherenkov radiation

- Charged particles can exceed the speed of light in media with indices of refraction $n > 1$.
- In this case they produce a cone of Cherenkov radiation with opening angle given by $\cos\theta_c = 1/(n\beta)$
- If $\beta < 1/n$ there is no radiation (threshold velocity)
- This type of radiation is useful in determining the particle type (e, μ , π , K or p) when combined with a measurement of particle momentum (measure p and v to determine m)

Multiple Coulomb scattering

- Charged particles scatter off the EM field (mostly of nuclei)
- The net effect of many small random scatters is a deflection in angle given by

$$\theta_0 = \frac{13.6 \text{ MeV}}{\beta c p} z \sqrt{x/X_0} \left[1 + 0.038 \ln(x/X_0) \right]$$

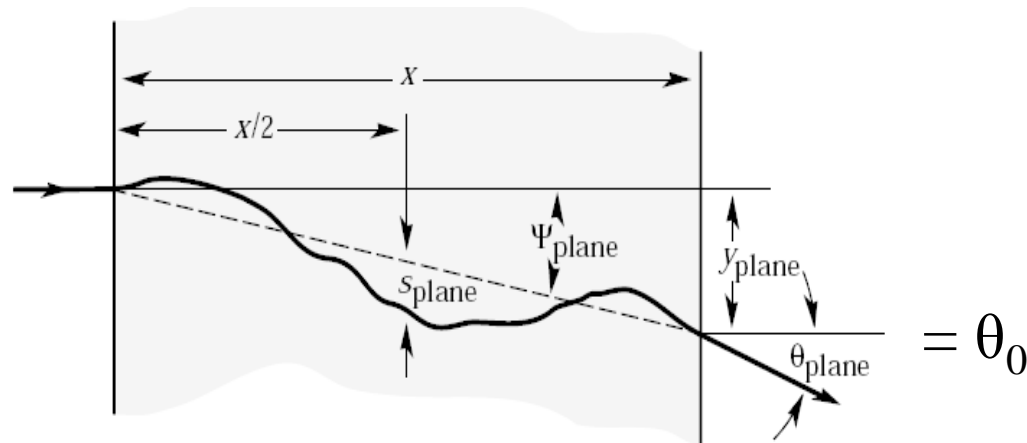
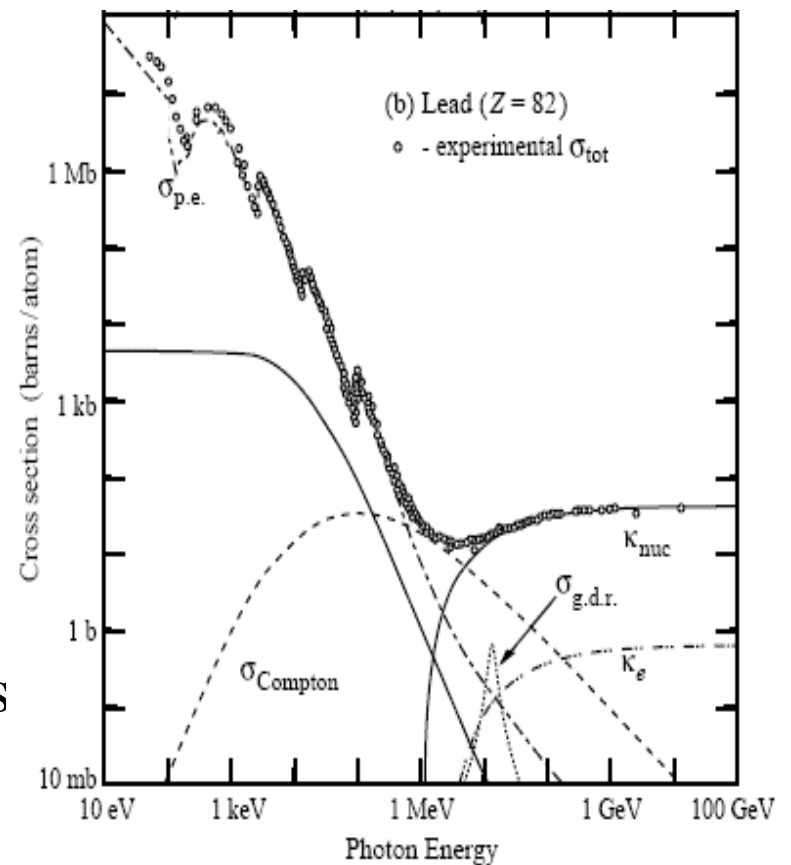


Figure 27.8: Quantities used to describe multiple Coulomb scattering. The particle is incident in the plane of the figure.

- These deflections limit the accuracy with which particle trajectories can be measured

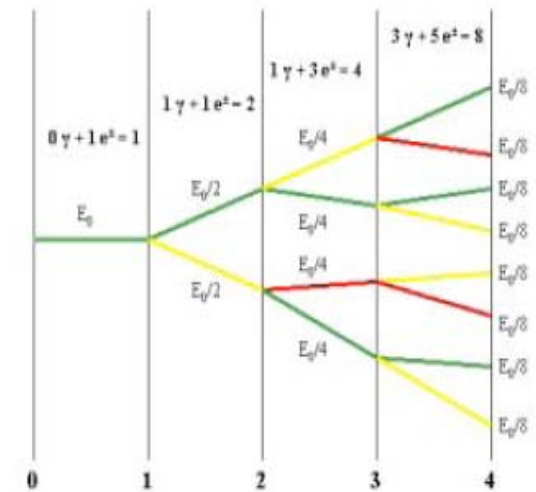
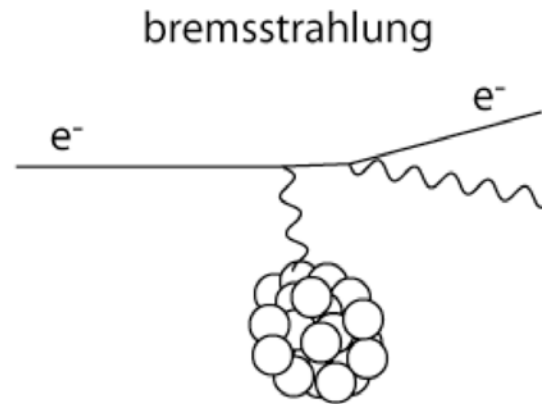
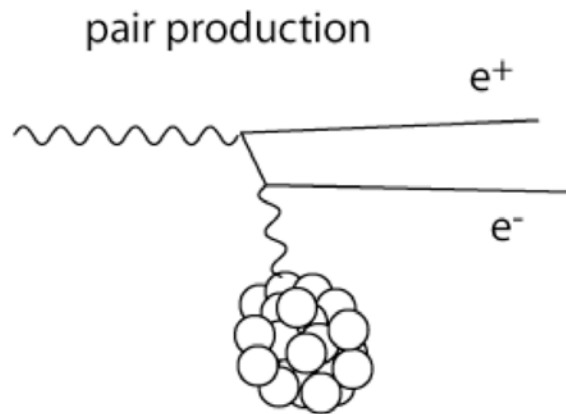
Photons

- Photons lose energy via
 - Photoelectric effect
 - Rayleigh scattering
 - Compton scattering
 - e^+e^- pair production
- The photons we're interested in measuring are always in the pair-production region
- The low-energy behavior is relevant to detector design due to electromagnetic showers, which produce many low-energy photons



Electromagnetic showers

- Both photons and electrons at high energy initiate a cascade of pair-production and bremsstrahlung that leads to a “shower” of particles



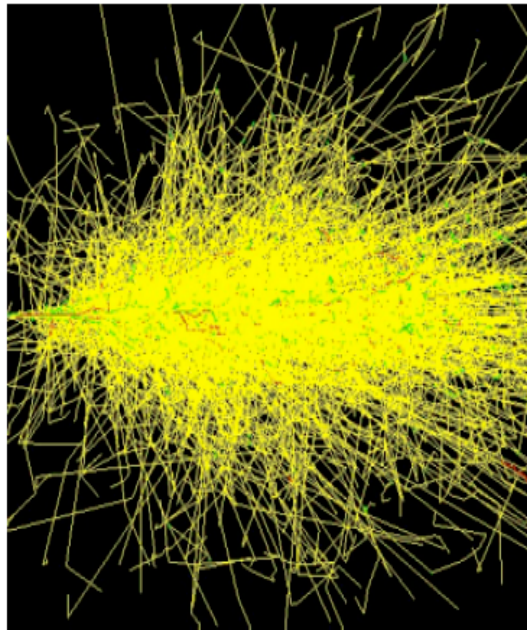
- Once the electrons and photons are at low energies they scatter (see previous slides) and result in a large number of low-energy photons

EM shower picture

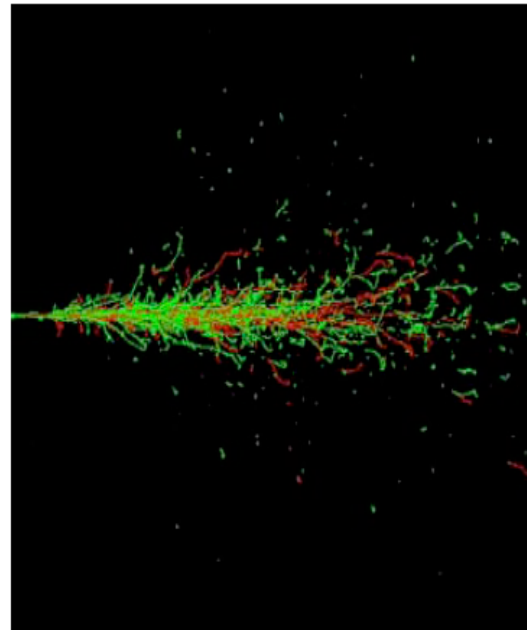
- Simulation shows development of a shower; play with it at <http://www2.slac.stanford.edu/vvc/egs/basicstool.html>
- The lateral width of the shower is determined by multiple scattering of the electrons in the medium (Moliere radius)

EGS simulation of 1 GeV electron shower in 15 cm of Cu (10 incident)

All particles are shown

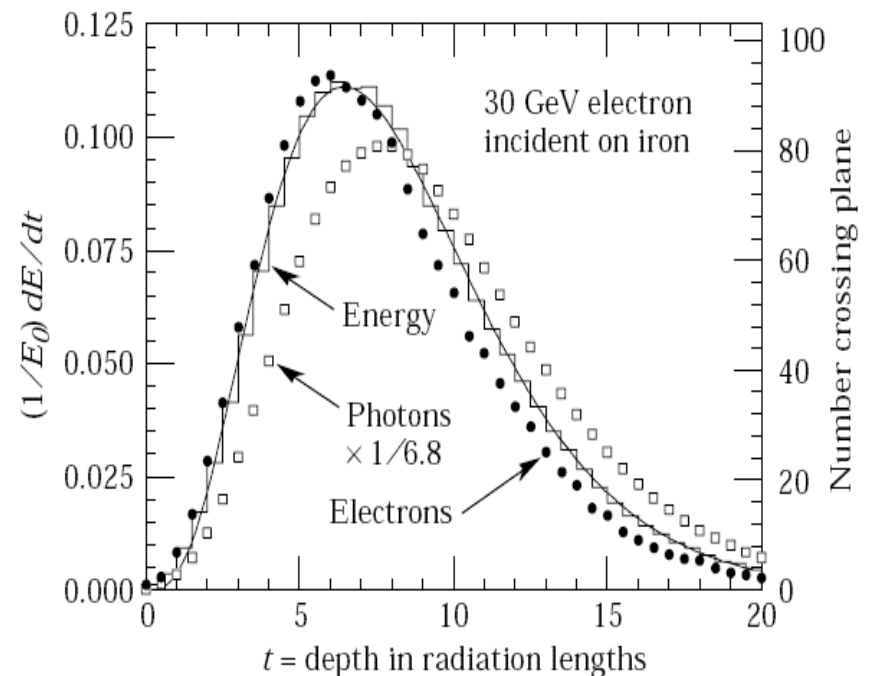


Only electrons are shown



EM shower characteristics

- The energy deposition varies with “depth” in the medium
- Depth measured in radiation lengths X_0
- Photon-induced showers start a bit later than electron-induced showers (initial pair production has to occur); effective radiation length for photons is $9/7 X_0$
- The energy of the photons at the end of the shower is proportional to the incident particle energy – a useful feature for building detectors



Hadrons

- All hadrons can interact strongly in matter
- Neutral hadrons (e.g. neutrons) interact *only* this way
- Hadrons create a cascade of particles (a hadronic shower), which produces mini-EM showers (from γ daughters of π^0) and loses *or* gains energy through nuclear interactions
- Hadronic showers are much less uniform and regular than EM showers
- The nuclear absorption length λ is analogous to the radiation length X_0 for EM showers

Detector strategies

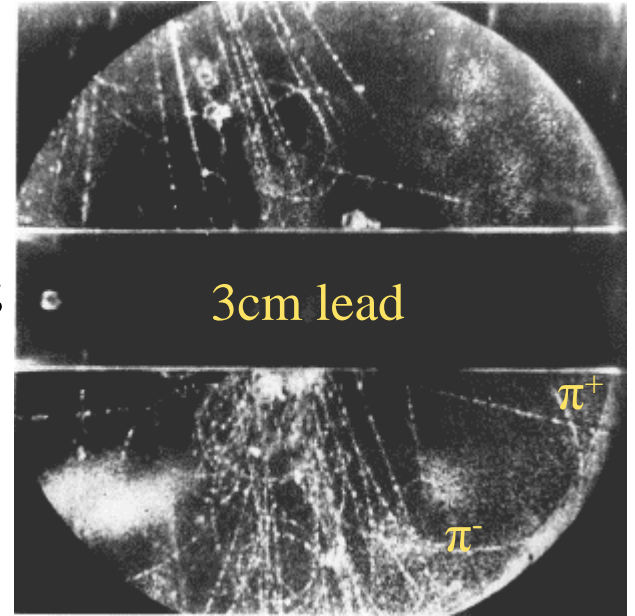
- “Non-destructive imaging” of charged particles
 - Use ionization energy loss, detect ionization trails
 - Use magnetic field to bend particles, determine momentum
- Determination of particle type
 - Measure speed in addition to momentum
 - Use presence/absence of Cherenkov light
- Absorption of photons / EM calorimetry
 - Use proportionality between incident energy and energy (or number) of photons in cascade
- Absorption of hadrons / hadronic calorimetry
 - Use rough proportionality (poorer resolution than for EM)
- Penetration of large amount of material → muons (or neutrinos)

Detectors of charged particles

- Magnetic field – dipole (for some fixed target detectors), solenoidal (for most colliding-beam experiments) or toroidal (large volume muon detection, e.g. ATLAS)
- Two main types:
 - Ionization based: gaseous mixture (usually including a noble gas, e.g. Ar or He)
 - Solid state, e.g. doped silicon, in which traversing charged particles create electron-hole pairs
- Both in widespread use.

Bubble and cloud chambers

- Early detectors: spark, cloud chambers
 - Spark chambers use HV and initiate a discharge along the ionization trail left by a particle (like lightning); poor spatial precision
 - Wilson cloud chamber uses supercooled vapor; ionization trail seeds phase transition to liquid (little drops form along the trajectory). Photographic readout.
 - Cycle time ~minutes



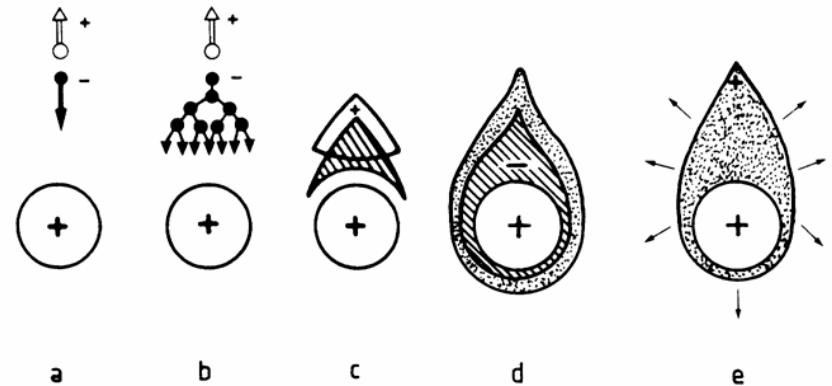
- Bubble chambers (Glaser)
 - Use superheated liquid; ionization trail seeds phase change to vapor
 - Good spatial resolution
 - Photographic readout (limits rate)
 - Pressure used to recondense liquid; cycle time ~1s

Scintillation detectors

- Ionizing radiation causes some materials (e.g. organic plastics and inorganic crystals) to “scintillate”, i.e. to release photons from the decay of molecules excited by the ionization.
- Scintillation light tends to be in the near UV; need “wavelength shifters” to facilitate optical readout
- Photomultiplier tubes record the generated light
- Scintillators are still in use; they are cheap (so large areas can be instrumented) and sensitive to the passage of a single charged particle.
- The fast decay time (few ns) of many scintillators makes them useful for triggering

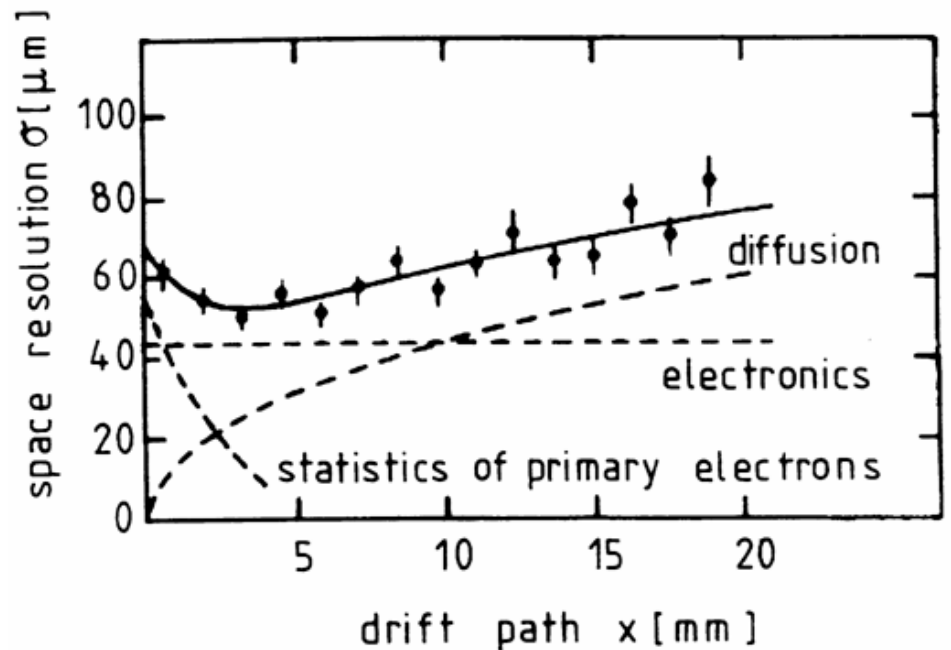
Wire chambers

- Multi-Wire Proportional Chambers (Charpak) allow ionization trails to be recorded electronically (no need for scanners!)
- Ionized electrons drift toward anode wires
- **E** field near anode $\sim \Delta V/r$
- Electrons are accelerated enough to cause an “avalanche” of ionization \rightarrow amplification of the signal (factor of 10^4 or so)
- Resolution determined by wire spacing
- Better idea: use arrival time to determine position (better resolution with fewer wires)



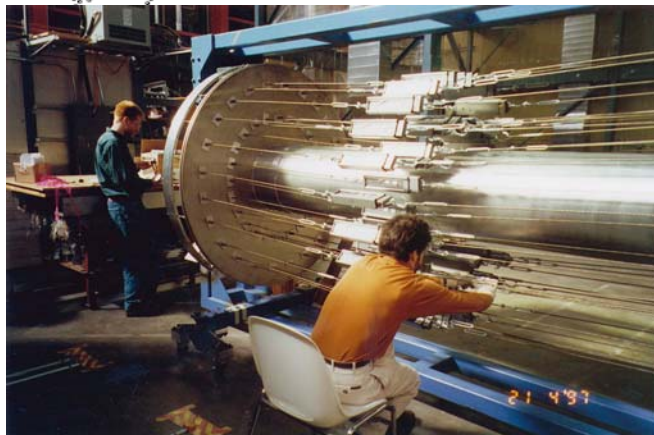
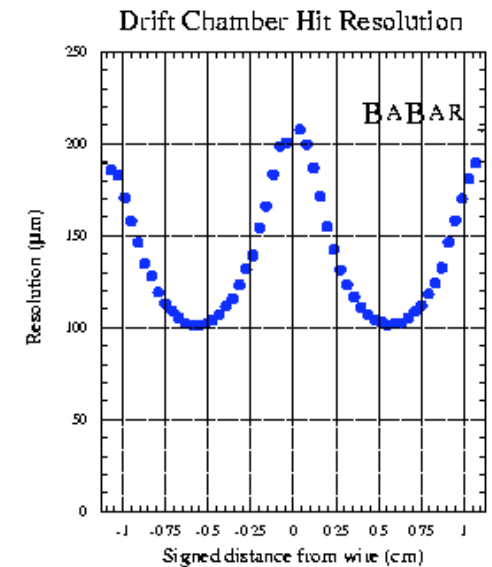
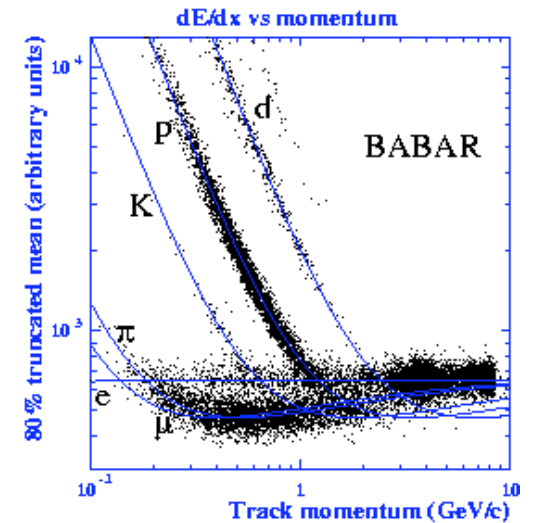
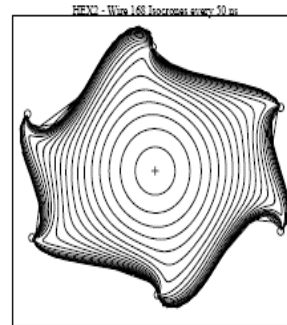
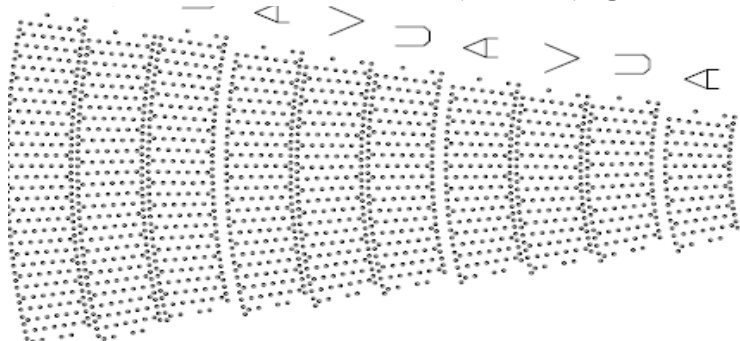
Drift chamber

- Ionization produced in a gas drifts in a modest electric field:
 - Drift speed is in the range of 10-100 $\mu\text{m}/\text{ns}$ (depends strongly on what gas is used)
 - Record the time required for the ionization to drift to the anode with \sim few ns timing precision, allowing \sim 100 μm measurements
 - Diffusion reduces resolution for long drift paths
 - Charge deposited (proportional to primary ionization) allows a simultaneous measurement of dE/dx energy loss \rightarrow particle identification



BaBar Drift Chamber

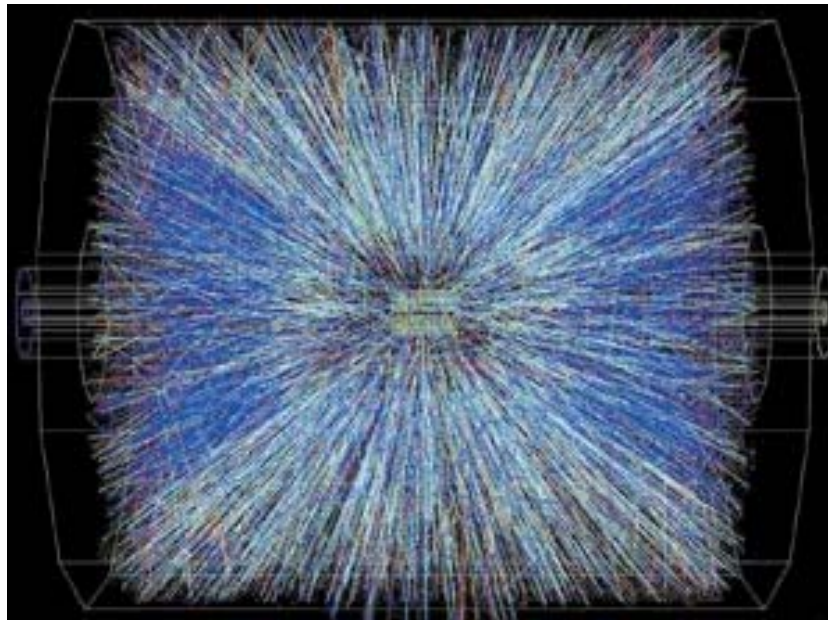
- Detector was strung at TRIUMF in 1997
- ~28000 wires; 7104 sense wires (20 μ m)
- 2.8m long, 0.2m to 0.8m in radius
- 4 axial and 6 stereo superlayers
- Helium-isobutane (80/20) gas mixture



Time projection chamber

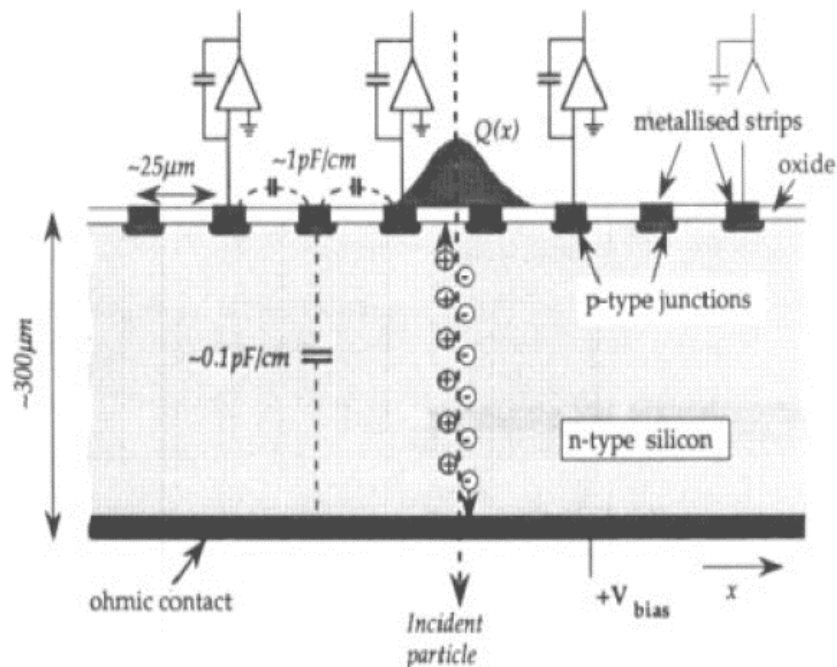
- Ionization trails drifted along axis of cylinder in *parallel E and B fields* \rightarrow electrons spiral along field lines, limits diffusion
- Long drifts (2m or more) are feasible
- Readout is entirely on endplates – no wires in the bulk volume
- Almost bubble-chamber—like pictures of trajectories:

STAR TPC at RHIC
(Au-Au collisions)



Silicon detectors

- Basic mechanism is electron-hole production along the path of a traversing charged particle in a reverse-biased region (i.e. where electrons and holes can't normally be found)
- Charge is collected on either strips or pixels bonded to the surfaces of the silicon wafer



Silicon detectors

- Excellent spatial resolution (5-50 μm)
 - Large number of electron-hole pairs produced ($\sim 10^4$ in 300 μm)
 - Diffusion is modest
 - Pitch (distance between strips) is small (25-200 μm); charge-sharing on nearby strips is used to determine track position (weighted average)
- Silicon detectors are relatively expensive, but large area ($\sim 100\text{m}^2$ of Si) detectors are being built for LHC
- Large Silicon detector arrays have up to 10 million readout channels $\rightarrow 10^7$ amplifiers, 10^7 ADCs, ... must use custom analog microchips for readout

Electromagnetic calorimeters

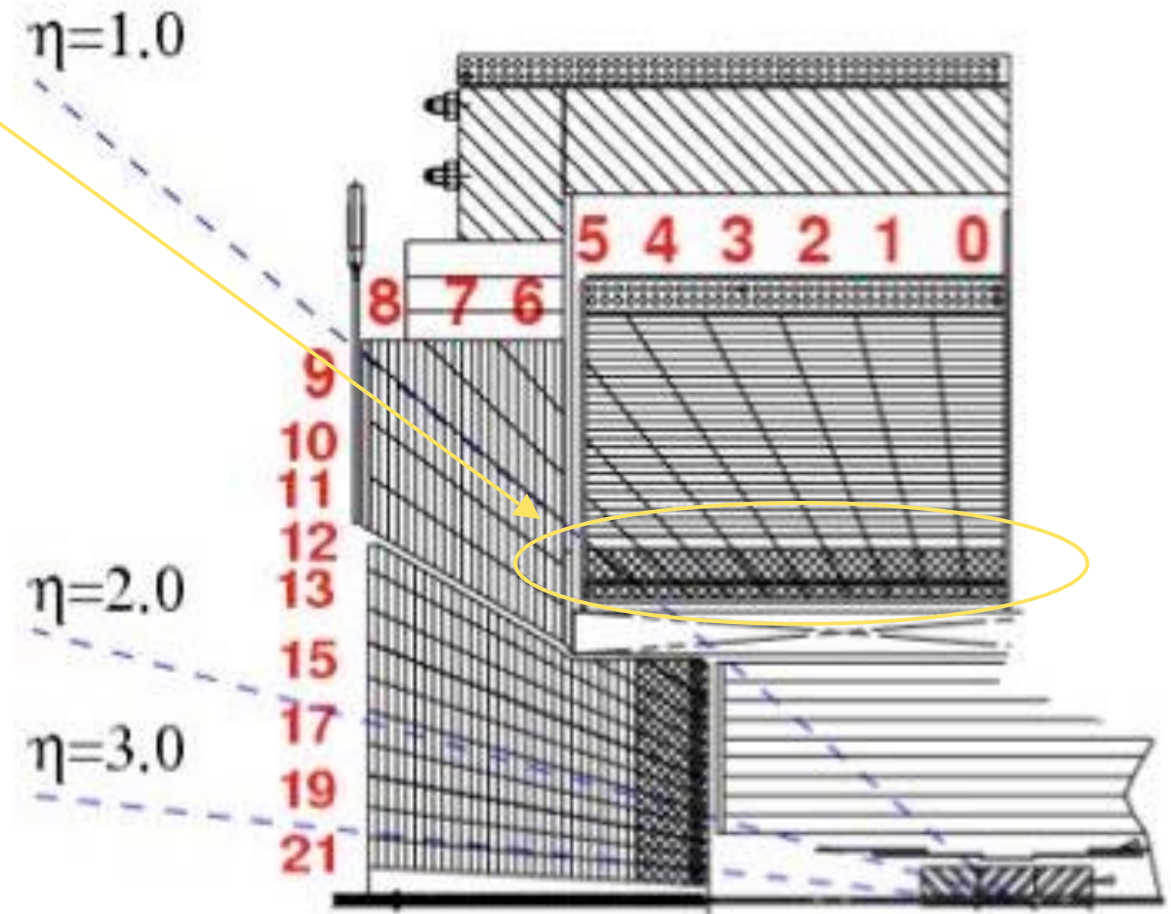
- Two main categories:
 - Sampling calorimeters, consisting of an absorber interspersed with active (readout) layers
 - Homogeneous calorimeters, where the entire volume is active
- Properties of interest
 - Energy resolution
 - Spatial resolution
 - Shower “shape”, i.e. spread and depth

Sampling calorimeters

- Stack of absorber (high Z material) and active medium (e.g. scintillator or wire chamber)
- Measure either the number or the energy of particles produced in the shower via a discrete number of samples
- Allows shower profile (depth) to be determined
- Energy resolution $\sigma_E \sim a/E^{1/2} + b$, where $a \sim 0.1$; dependence on $E^{1/2}$ comes from measuring $N \pm \sqrt{N}$ photons (since E is proportional to N)

CDF sampling calorimeter

- calorimeter is arranged in projective “towers” that point to the interaction region
- Electromagnetic and hadronic sections are both sampling



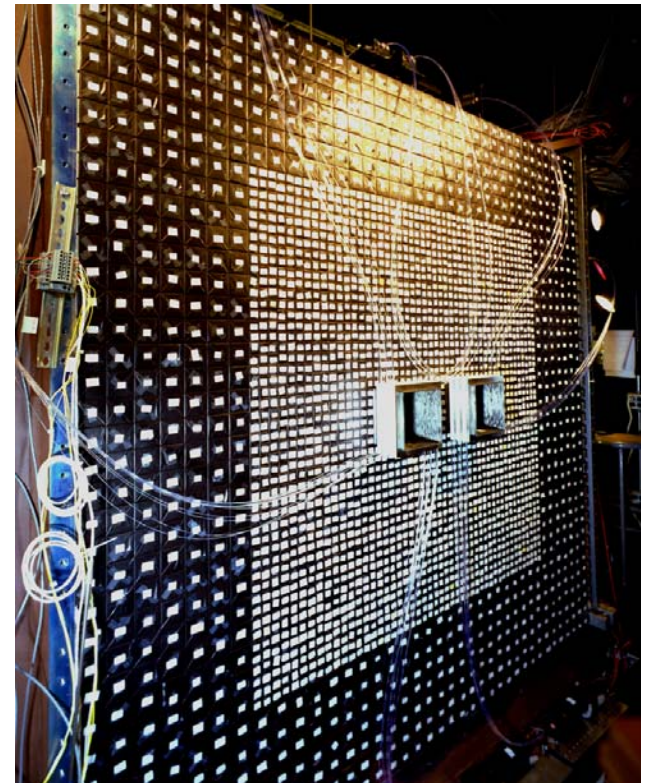
Crystal calorimeters

- Shower develops and is read out in the same medium; no sampling fluctuations to reduce resolution
- Light is produced by scintillation or Cherenkov radiation
- Longitudinal segmentation is more difficult than for sampling
- Good energy resolution: $\sigma_E \sim a/E^{1/4} + b$, where $a \sim 0.03$ and E is measured in GeV



CsI crystal
being wrapped

KTeV calorimeter

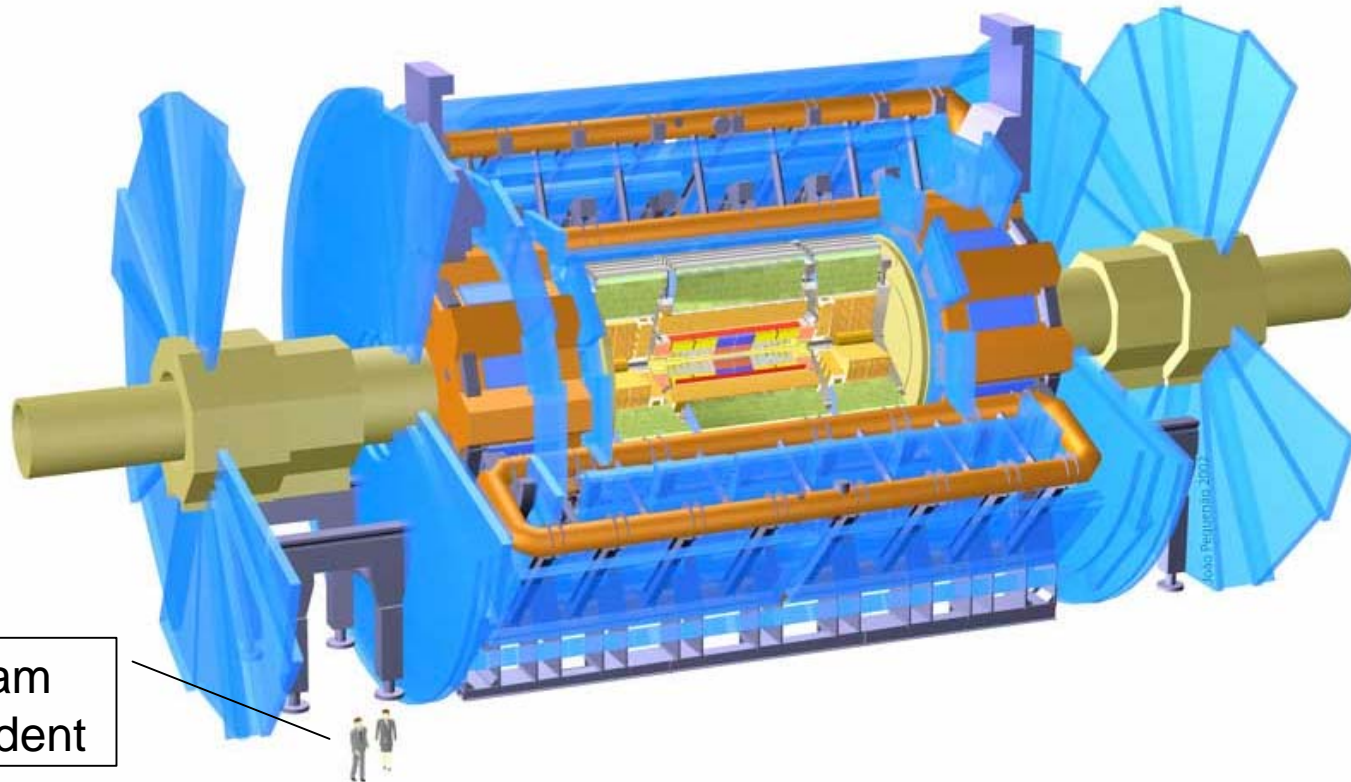


Hadronic calorimetry

- Very different response to hadrons (e.g. π^+) and γ (which come predominantly from π^0 decay)
- Sampling calorimeters tend to be used (e.g. Cu and liquid Argon for ATLAS)
- Try to design calorimeter to equalize the response to these components of the hadronic shower (compensating)
- Resolution poorer than for electromagnetic calorimetry; typical values $\sigma_E \sim a/E^{1/2}$ with $a \sim 0.4-1.0$

Muon detectors

- Consist of wire chambers (or scintillators) separated from particle source by substantial shielding
- Sometimes placed in magnetic field to allow independent measurement of muon momentum (e.g. ATLAS)



Sue and Sam
graduate student

Triggering, readout, reconstruction

- Particles from interesting interactions leave signals in the detector, but so do particles from common and uninteresting processes, from electronic noise, etc.
- Reading the information (times and pulse heights for each channel) is impractical; it takes too long, and during the process the detector cannot record new events
- Triggering attempts to decide which events are worth reading out based on a subset of information obtained quickly; reduction factors are often 10^6 or more
- Once a set of events is read out and stored, sophisticated algorithms reconstruct particles from the channels with signals and determine momenta, positions and energies with precision

Modern colliding-beam detectors

- Layered concept – first tracking, then particle identification, then calorimetry, then muon detection
- Many examples – size dictated by maximum particle momentum
- Variety of technologies – dictated by performance, cost, accumulated expertise
- Huge undertaking – requires hundreds of scientists working together

BaBar detector

- covers $\Omega = 0.90 \times 4\pi$ in CM frame
- Asymmetric
- Si strips at vertex
- Drift chamber
- Ring Imaging Cherenkov detector
- CsI(Tl) calorimeter
- Muon system

